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Dust Temperature Fluctuations and Surface Processes: Impact on H₂ Emission Lines in PDRs.

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H₂ emission lines are a powerful tool to study PDRs, as they give informations about the conditions near the H/H₂ transition. However, the modeling of H₂ excitation state and the H/H₂ transition – which then controls the apparition of other species, e.g. the CO transition – requires a careful understanding of several physico-chemical processes. Two remain problematic in models : the H₂ formation process, which controls the position of the H/H₂ transition and can induce formation pumping, and the slow ortho-para conversion processes, with a competition between reactive collisions in the gas (with H⁺ and H) and thermalization on dust surfaces.

ISO and Spitzer observations of H₂ emission lines provide constraints for the models, and show very efficient H₂ formation in PDRs (Habart et al., 2011). The introduction of detailed H₂ formation mechanisms in the Meudon PDR Code has led to an improved formation efficiency at the edge, in better agreement with the observations (Le Bourlot et al., 2012), but discrepancies remain in low-UV-flux PDRs, related to the ortho-para conversion processes.

Dust grains are the support of both the formation mechanism and an efficient ortho-para conversion process. Small grains are the largest contributors as they account for most of the total dust surface. But small grains undergo temperature fluctuations caused by the absorption of individual UV photons. As the surface processes are very sensitive to the dust temperature, these fluctuations need to be taken into account, but are problematic for both stationnary and time-dependent PDR models (due to the short timescales involved).

I have thus developed a statistical formalism based on master equations to compute the average efficiency of surface processes perturbed by fluctuations of the grain temperature. I present here the application of this method to H₂ formation (Bron et al., 2014) and ortho-para conversion, and show the strong impact of fluctuations in PDRs. The integration of this method with the Meudon PDR Code then allows us to directly compare to H₂ observations in PDR.

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